# **Preparing for LHC Physics**

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LHC accelerator and detectors are being installed — plan is for first collisions in "summer of 2007".

Commissioning will take time: initially just 43/2808 bunches. Hope for  $\lesssim 100 \,\mathrm{pb^{-1}}$  in 2007. But need to start getting ready now....

#### Topics:

- NLO QCD corrections to  $b\bar{b}H$  [Sally Dawson, Chris Jackson]
- NNLO QCD corrections to  $Z \to \ell^+ \ell^-$  [Bill Kilgore].
- SUSY phenomenology separate talk [Tadas Krupovnickas].
- Preparations for ATLAS physics [F.P.; see also Kyle Cranmer]



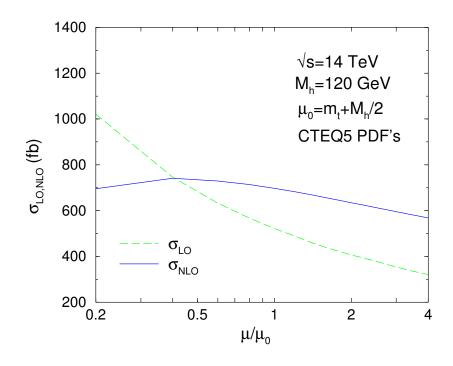
## **NLO** *bbH* Corrections

Cross section for  $b\bar{b}H/A$  is enhanced for  $\tan\beta \gg 1$ . Decay  $H/A \to \tau^+\tau^-$  also enhanced. Not early physics but potentially important.

Calculation [Dawson, Jackson, Reina, Wackeroth] extends previous work on  $t\bar{t}H$ . Compute one-loop corrections to  $gg \to b\bar{b}H$  (massive pentagons) and corresponding real emission processes. Essential for double b tag.

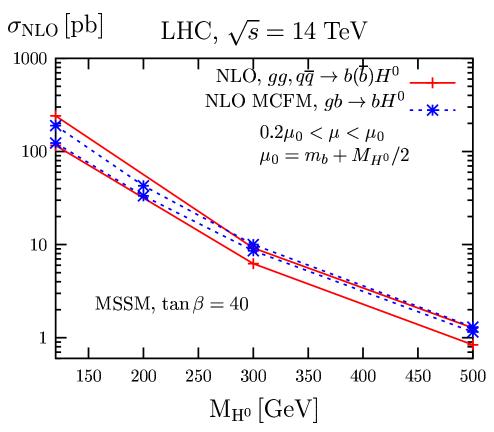
Much reduced scale dependence – NLO result looks reliable.

Note choice of relatively low scale improves convergence. Confirms previous conjecture.





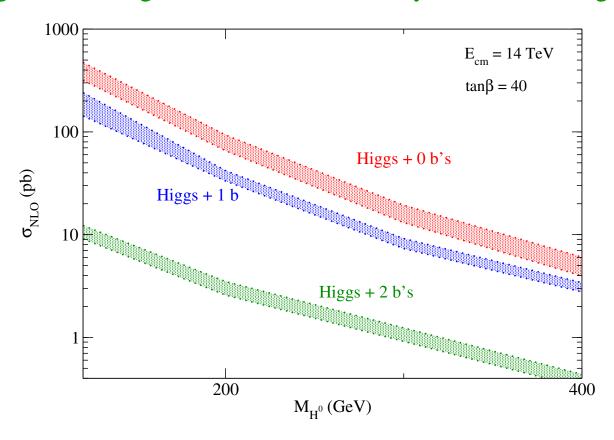
For single b tag can treat process as  $gb \to Hb$ . Much simpler calculation available in MCFM [Campbell, Ellis]. Require  $p_{T,b} > 20 \,\text{GeV}$ ,  $\eta_b < 2.5$  and compare:



Very good agreement.

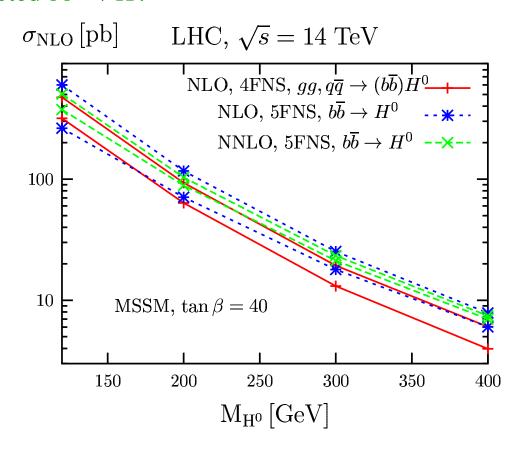


Requiring second b tag reduces cross section by an order of magnitude:



Double tag is not used in ATLAS analysis. Might be useful cross check.

Can also compare inclusive rate without b tag with NNLO calculation [Harlander, Kilgore]. Good agreement between NLO-corrected  $gg \to b\bar{b}H$  and NNLO-corrected  $b\bar{b} \to H$ :



Conclusion: cross section is under good control.



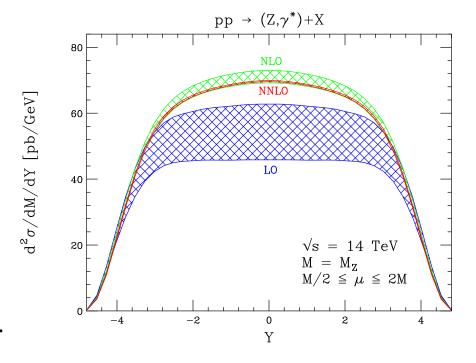
## **NNLO** $W \to \ell v$ and $Z \to \ell^+ \ell^-$ Corrections

 $W \to \ell v$  and  $Z \to \ell^+ \ell^-$  will be early LHC physics: expect  $\sim 10^5 Z$  and  $\sim 10^6 W$  events for  $100 \, \mathrm{pb}^{-1}$ . Crucial for calibration and for luminosity determination — traditional methods difficult at LHC energies.

Integrated cross section was first NNLO calculation for *pp*.

Extended to  $d\sigma/dy$  with clever trick [Anastasiou, et al.]. But not enough to calculate acceptance.

Need full production/decay distribution to determine acceptance and hence luminosity.





Calculation [Bill Kilgore] uses extension of NLO subtraction:

$$\sigma_{\text{NLO}} = \int_{n+1} d\sigma_R + \int_n d\sigma_V 
= \int_{n+1} \left[ d\sigma_R - d\sigma_A \right] + \int_{n+1} d\sigma_A + \int_n d\sigma_V 
= \int_{n+1} \left[ d\sigma_R |_{\epsilon=0} - d\sigma_A |_{\epsilon=0} \right] + \left[ \int_{n+1} d\sigma_A + \int_n d\sigma_V \right]_{\epsilon=0}$$

Compute first term numerically and second analytically.

For NNLO have double real, real-virtual, and 2-loop virtual with multiple overlapping subtractions. Exploit fact that total is finite.

Analytic part is finished: have complete expression for  $d\sigma_A$  and for analytic integral. Latter can be written as  $L^{\mu\nu}H_{\mu\nu}$ , where  $L^{\mu\nu}$  is elementary lepton tensor and  $H_{\mu\nu}$  is expressed as combination of dilogarithms and elementary functions. Result for (dominant)  $q\bar{q} \to ZX$  part:



Full analytic result exists, but expression is so large that single-page Postscript file was too slow to display(!).

Result gives  $d\sigma/d^3p_{\ell_1}d^3p_{\ell_2}$ . Hence can integrate with arbitrary lepton cuts. Of course result is singular if one forces, e.g.,  $\vec{p}_{T,\ell_1} + \vec{p}_{T,\ell_2} = 0$ . Need NNLO resummation to handle this.

Remaining task is to implement numerical calculation. Not trivial: must deal with numerical cancellation of  $d\sigma_R - d\sigma_A$ .

Expect results well before first  $Z \to \ell^+ \ell^-$  is observed at LHC.

Calculation is significant advance in state of the NNLO art.

But Z production is somewhat special: double singular regions only for  $p_{T,Z} = 0$ . Need more new techniques for, e.g., NNLO jet cross section.

And backgrounds for new physics typically involve many partons....



## **ATLAS Preparations for Physics**

Most previous ATLAS physics studies have considered 10–100 fb<sup>-1</sup>. Now beginning Computing System Commissioning (CSC). Main goals:

- Simulate "as-built" detector: e.g., mis-alignment of beam, solenoid, and calorimeter by few mm.
- Analysis strategy for first  $100 \,\mathrm{pb}^{-1}$ .

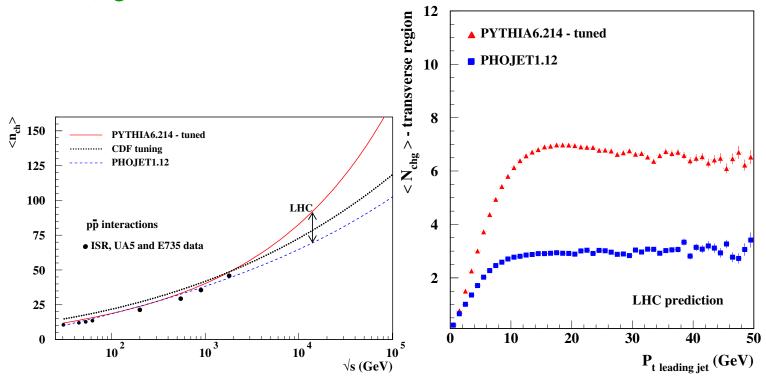
Initial luminosity unclear, but will have  $100 \,\mathrm{pb}^{-1}$  before  $100 \,\mathrm{fb}^{-1}$ .

Even  $100 \,\mathrm{pb^{-1}}$  is not small: gives about  $10^5 \,Z \to \ell\ell$ , about  $10^5 \,t\bar{t}$  and perhaps  $10^3 \,\mathrm{SUSY}$  events. Major discovery is possible!

Plan to simulate with GEANT (15–30 m/event) and to reconstruct (1 m/event) about 10<sup>7</sup> physics events. Mainly Standard Model processes/backgrounds. Huge but necessary effort....



Even tiny data sample would be useful: do not understand soft physics. Extrapolations of average multiplicity (left) and underlying multiplicity in jet events (right) [Arthur Moraes]:



Hope for much more: detector calibration, background measurement, even search for new physics. Will discuss each briefly....

**Calibration:** ATLAS calorimeter optimized for EM resolution, driven by  $h \rightarrow \gamma \gamma$ . Hence hadronic response has e/h > 1.

Current reconstruction assumes correct EM calibration and applies H1-style correction: EM showers are dense, so  $E_{\text{true}} \approx E$ ; hadronic showers are difuse, so  $E_{\text{true}} > E$ .

Existing calibration OK to few percent based on comparing reconstructed jets with Monte Carlo truth.

OK for discovery but not for precision measurements. Goal is  $\lesssim 1\%$ .

Hadronic shower simulations are not reliable. Need to test calibration *in* situ using, e.g.,  $p_T$  balance between  $\gamma$  or  $Z \rightarrow ee$  and jets, mass for  $t \rightarrow q\bar{q}b$ , .... Work on this is just starting.

**Backgrounds:** For narrow peak (e.g., Z' or KK resonance) can measure backgrounds from sidebands.

SUSY is perhaps most likely early discovery physics. Cross section is  $\lesssim 10 \, \mathrm{pb}$ , so must rely on inclusive signatures:

- $E_T$  plus multijets.
- $E_T$  plus hard dijets (e.g., from  $q_R \to \tilde{\chi}_1^0 q$ ).
- Possibly dileptons plus  $\mathbb{E}_T$ .

Existing analyses have estimated backgrounds using parton shower Monte Carlo plus fast detector simulation. Neither very reliable. Hence make hard cuts  $\Rightarrow$  negligible SM background.

Cannot afford to do this with limited statistics.

For first time have enough fully simulated events to study background. Using multi-parton matrix element generators (ALPGEN, SHERPA).

But ALPGEN, SHERPA,... are leading order in QCD — have large scale dependence. NLO calculations of SUSY backgrounds (e.g., Z+4jets) not available in forseeable future.

Must measure backgrounds from data.

For  $Z \to v\bar{v} + n$ jets, can simply measre  $Z \to \ell^+\ell^- + n$ jets with small  $E_T$ .  $t\bar{t}$  is intermediate in difficulty [Dan Tovey].

QCD multijets are probably most difficult background to understand. Contributions both from heavy flavor  $(b, c \rightarrow vX)$  and from mismeasured jets (cracks, shower leakage, ...).

High order in  $\alpha_s$ , so NLO calculation impossible.

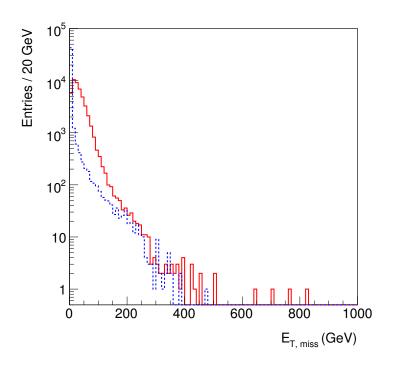
Must measure samples minimizing SUSY "background":

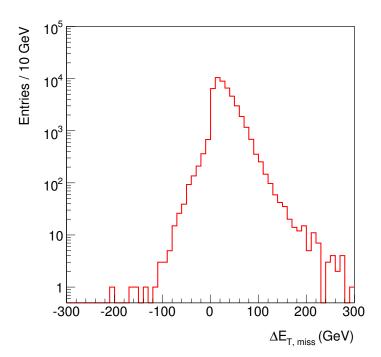
- Multiple jets with small  $\mathbb{E}_T$ ;
- Coplanar dijets (and perhaps  $\gamma$ +jet) with large  $\mathbb{E}_T$  from crack, leakage, etc.



Then assume jet mismeasurement factorizes: combine mismeasurement probability with multijet cross section. Not easy, but for first time will have sufficient data to study problem.

Simulation of dijets of with  $560 < E_T < 1120\,\text{GeV}$  for  $E_T$  (left) and  $\Delta E_T$  (right) looks OK for  $\Delta E_T \gtrsim 100\,\text{GeV}$ :





Need to verify this with real data.



#### **Conclusion**

Importance of TeV scale has been understood for at least 25 years. LHC is about to give us first data.

And yes, we need more postdocs.

